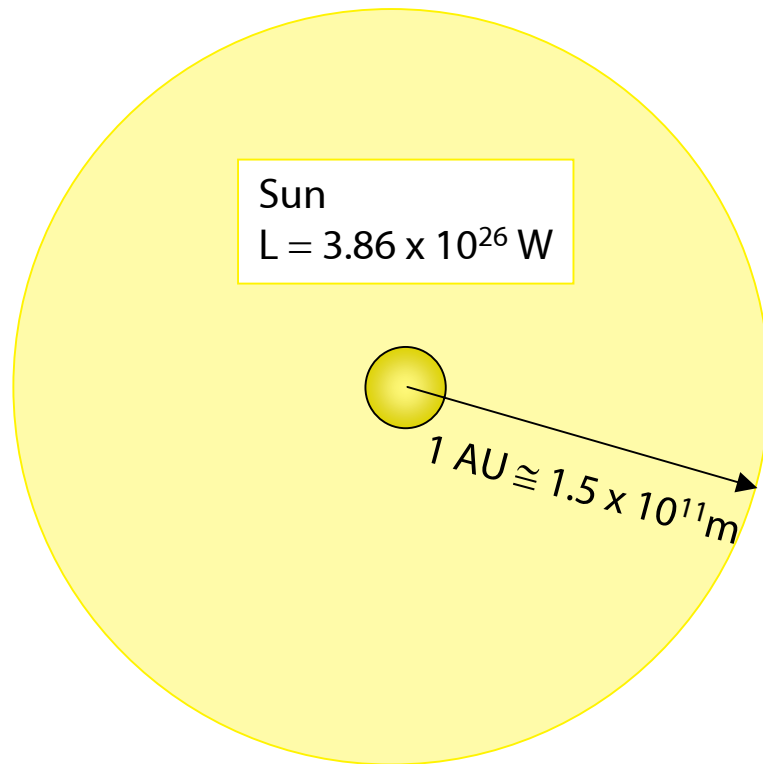


UNIT 2: Solar Geometry

- ▶ Solar constant
- ▶ Extraterrestrial radiation
- ▶ Sun-earth astronomy
- ▶ Examples

Solar Constant



Area of a sphere = $4\pi r^2$

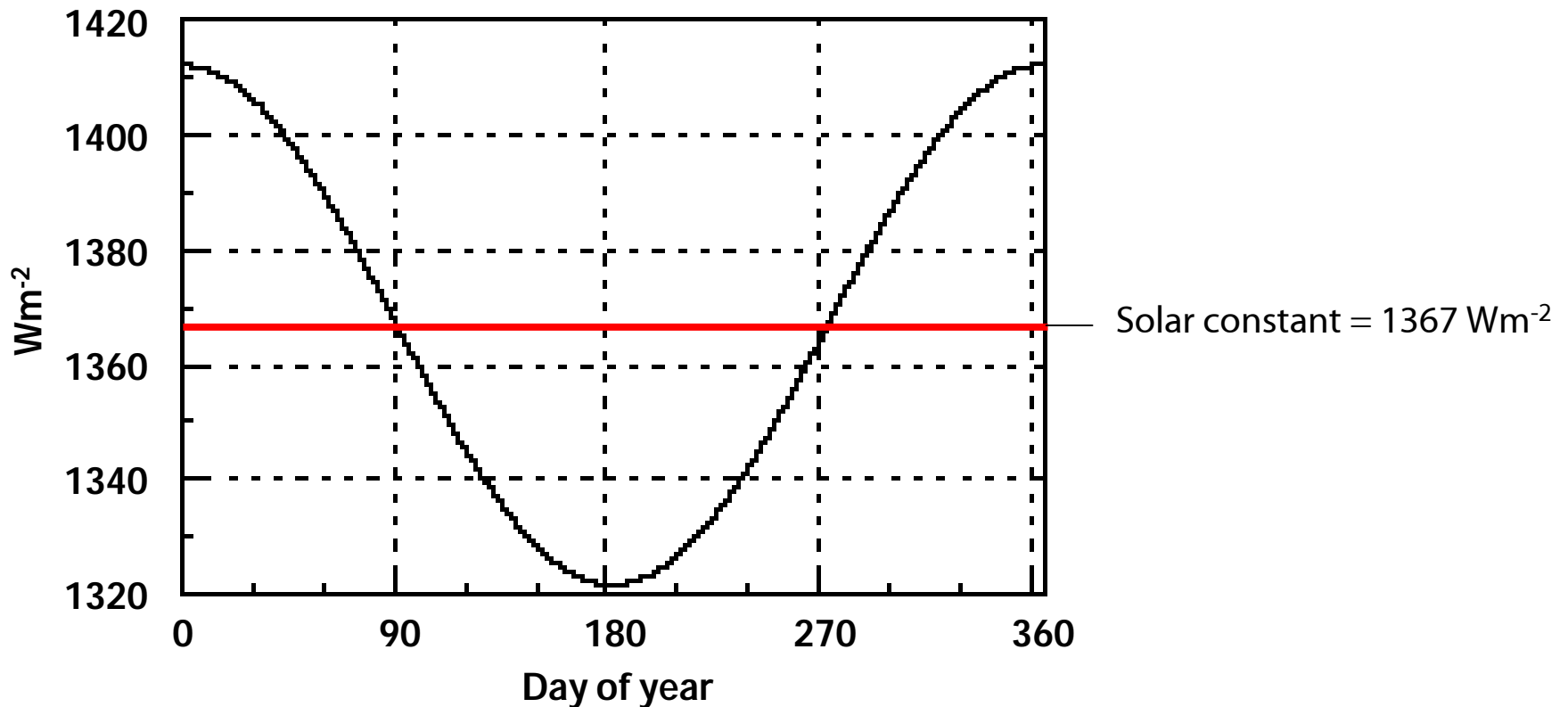
Conservation of energy requires that the total energy flux coming out of the sun must also pass through a sphere at 1 AU.

The energy flux density at 1 AU is

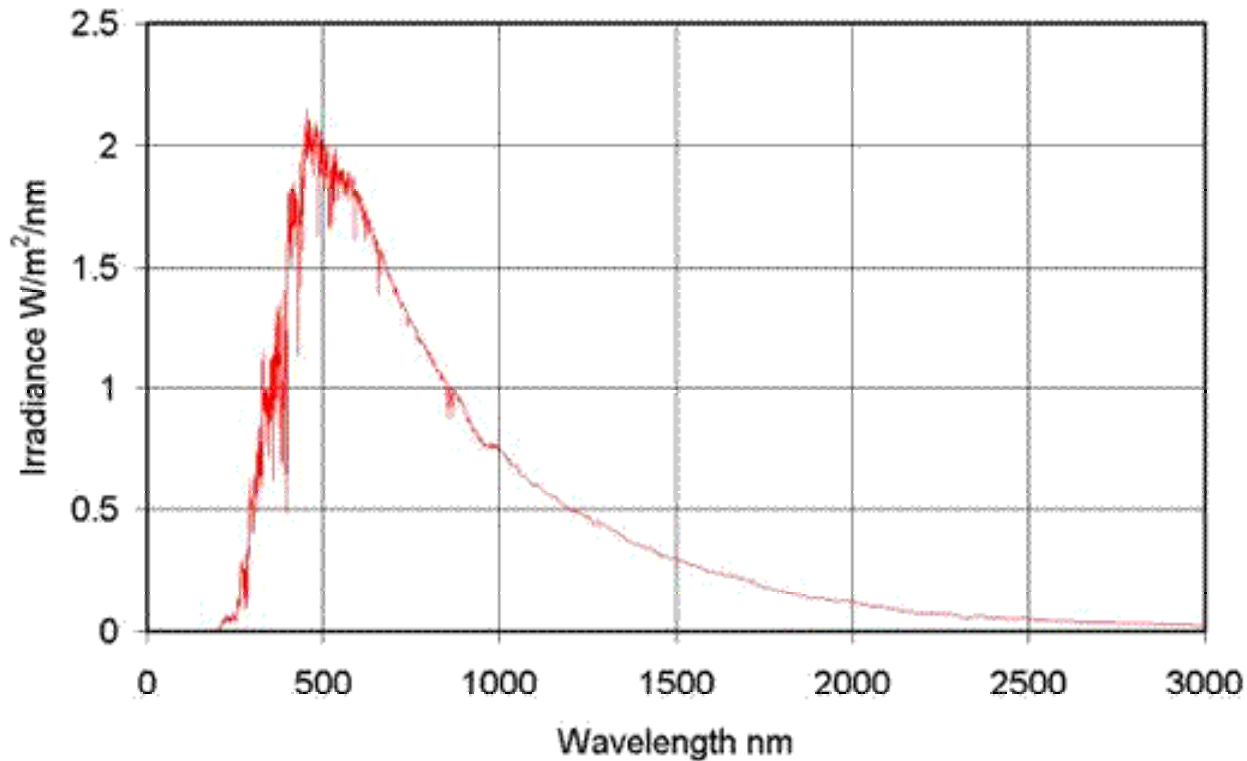
$$\frac{L}{4\pi r^2} = 1367 \text{ Wm}^{-2}.$$

This is the **Solar Constant**.

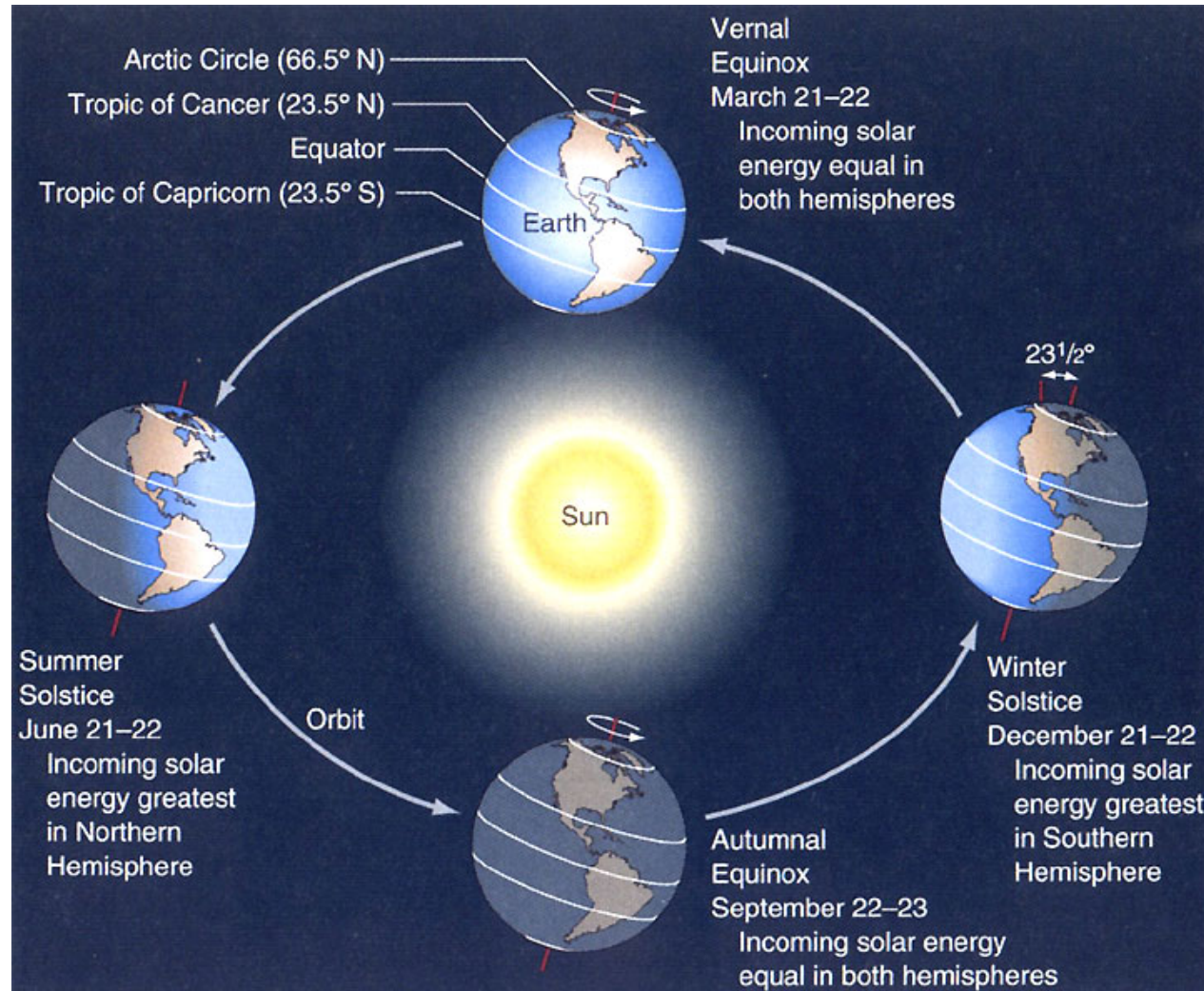
Extraterrestrial Radiation



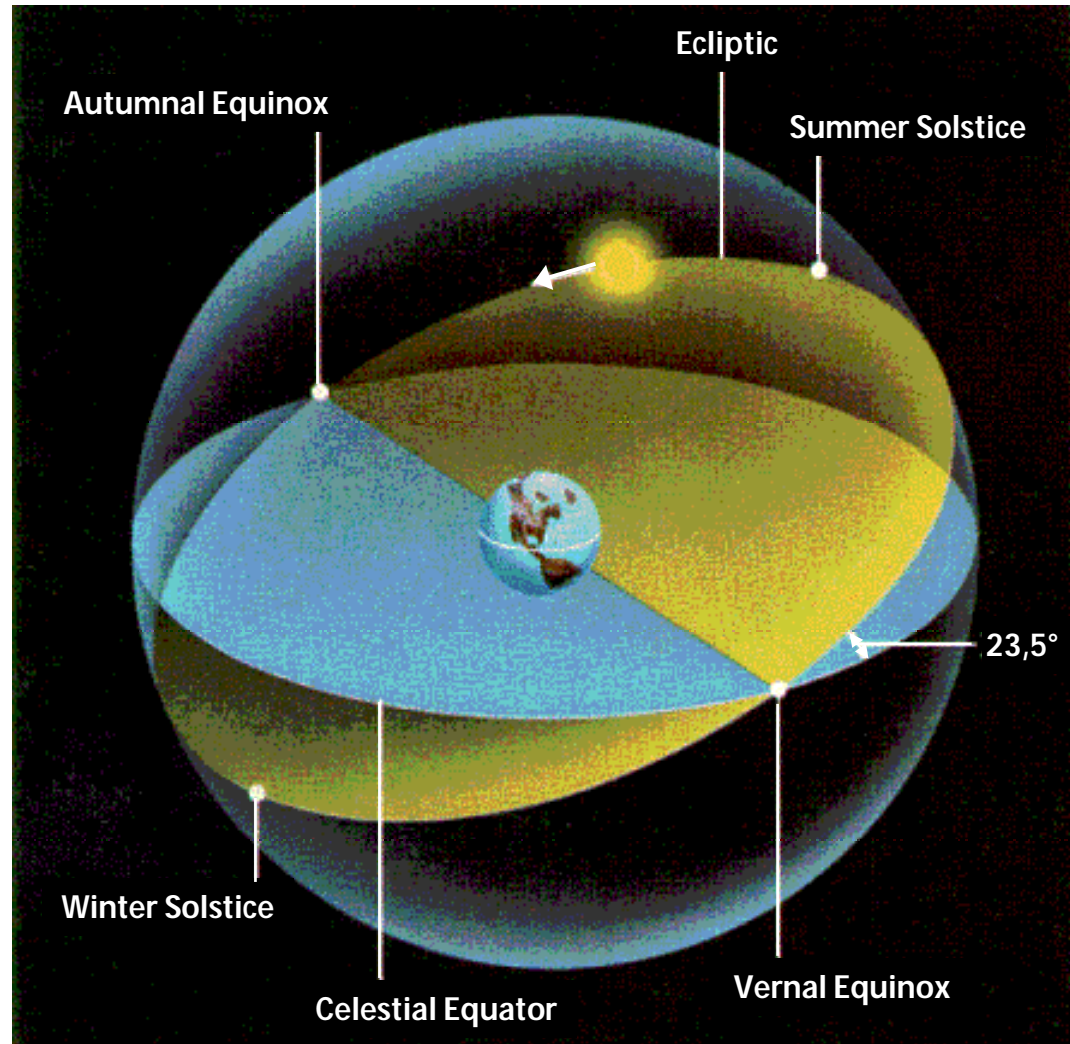
Extraterrestrial Solar Spectrum



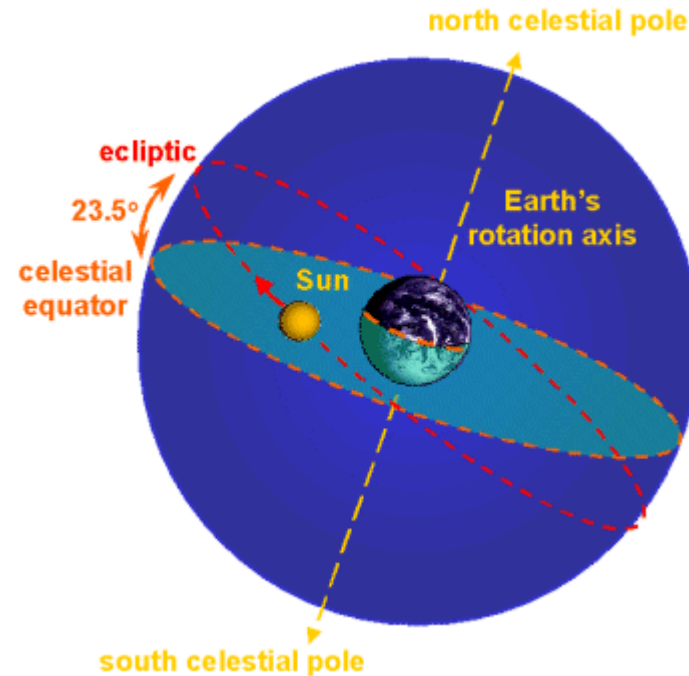
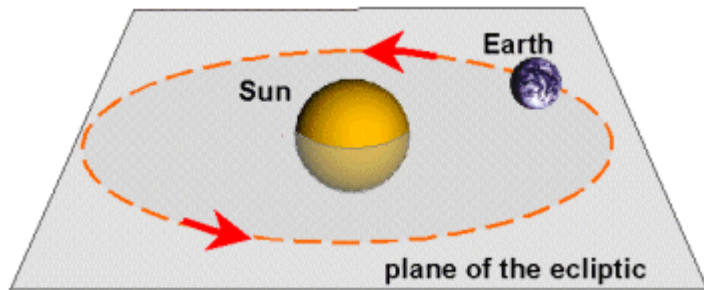
Ecliptic



Ecliptic



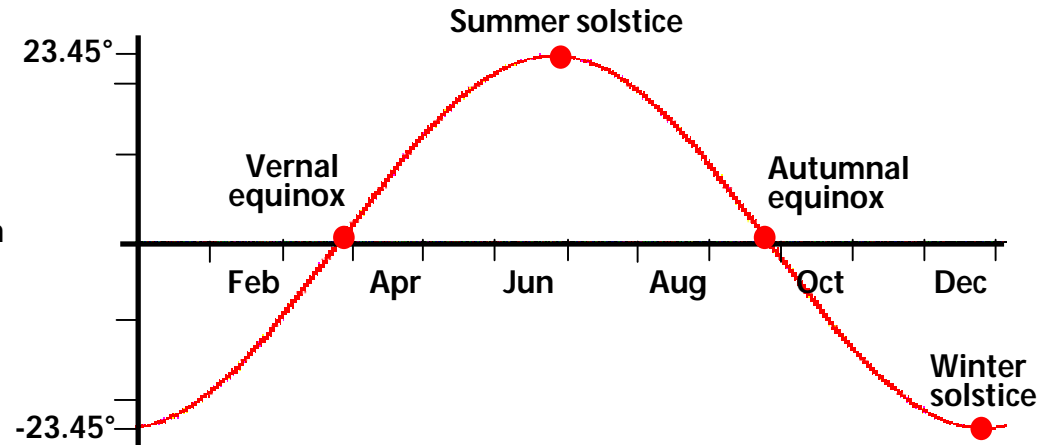
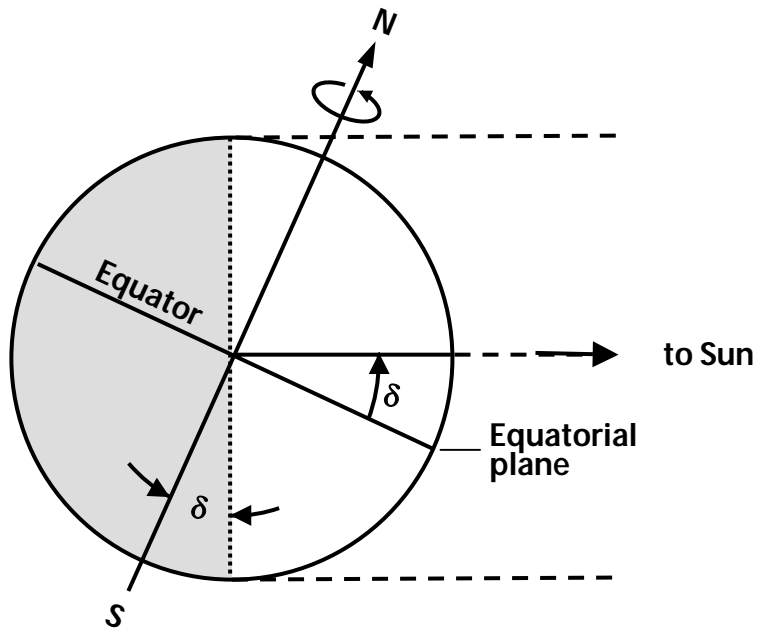
Ecliptic



The ecliptic is the region of sky (region of the [celestial sphere](#)) through which the Sun appears to move over the course of a year. This apparent motion is caused by the Earth's orbit around the Sun, so the ecliptic corresponds to the projection of the Earth's orbital plane on the [celestial sphere](#). For this reason, the Earth's orbital plane is sometimes called the plane of the ecliptic.

Due to the tilt of the Earth's rotation [axis](#) with respect to its orbital plane, there is an angle of 23.5° between the ecliptic and the celestial equator.

Solar Declination



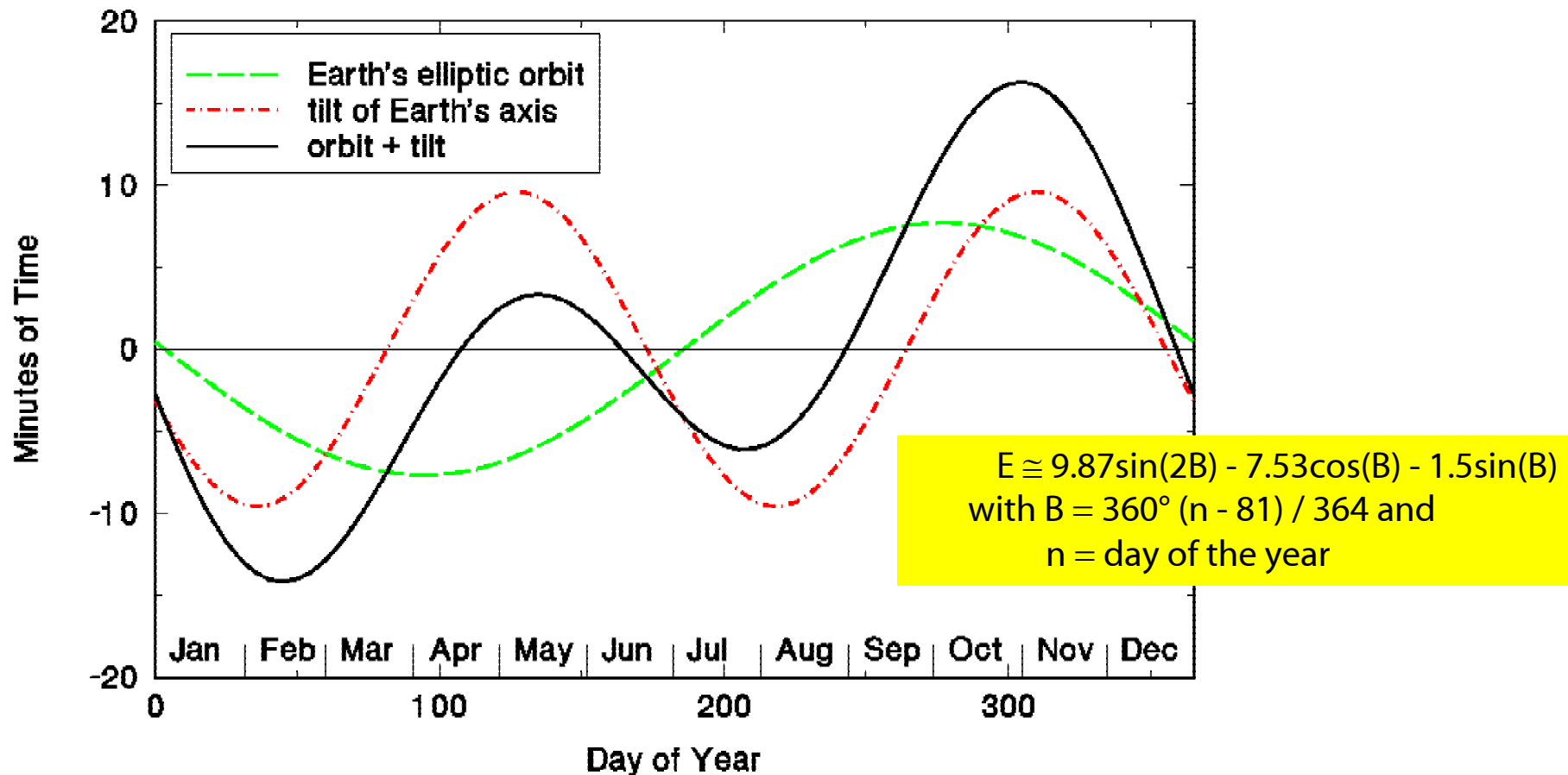
Declination angle $\delta = 23.45^\circ$

Variation of the declination angle:

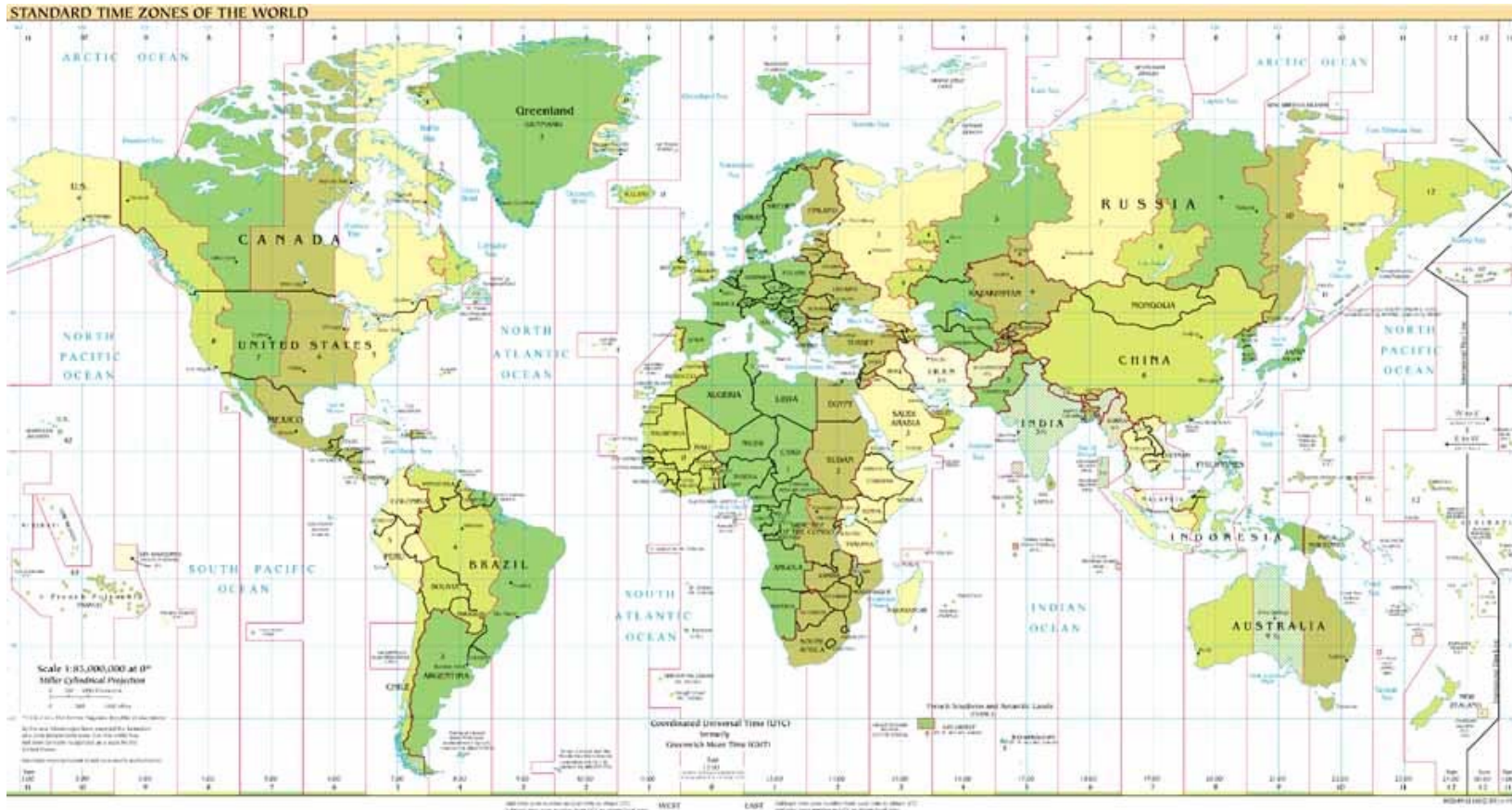
$$\delta \cong 23.45 * \sin [360 / 365 * (284 + n)]$$

with $n = \text{day of the year}$

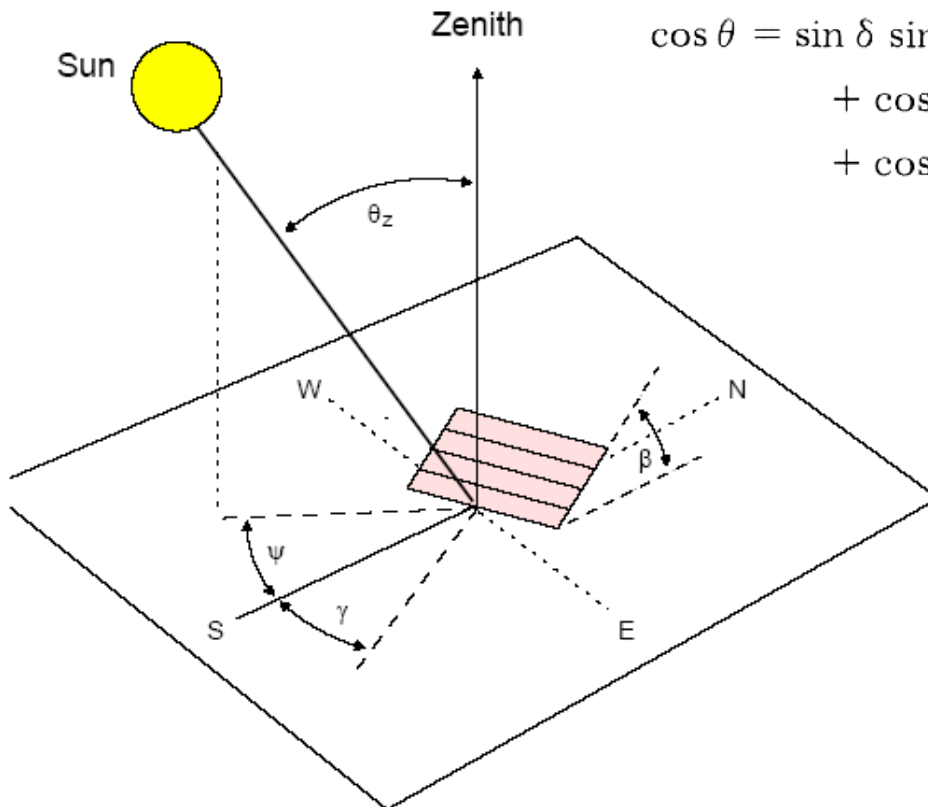
Equation of Time



Time Zones



Solar Geometry: Angle of incidence θ



$$\begin{aligned} \cos \theta = & \sin \delta \sin \phi \cos s - \sin \delta \cos \phi \sin s \cos \gamma \\ & + \cos \delta \cos \phi \cos s \cos \omega + \cos \delta \sin \phi \sin s \cos \gamma \cos \omega \\ & + \cos \delta \sin s \sin \gamma \sin \omega \end{aligned}$$

with:

latitude ϕ

solar declination δ

hour angle ω

slope s

surface azimuth γ